

# Thermal Properties of Crystal Lattices



**Peter Debye**  
(1884 – 1966)

The motion of a harmonic crystal can be described by a set of decoupled harmonic oscillators


$$H = \frac{1}{2} \sum_{\mathbf{k}, s} |\mathbf{P}_s(\mathbf{k})|^2 + \omega_s^2(\mathbf{k}) |\mathbf{Q}_s(\mathbf{k})|^2$$

The occupancy of a given mode is

$$\langle n_s(\mathbf{k}) \rangle = \frac{1}{e^{\beta\omega_s(\mathbf{k})} - 1}$$

- ➡ Thermodynamic properties of the ionic lattice
- ➡ long-range order in the presence of lattice vibrations

$$\begin{aligned} \frac{d}{dt} \langle \mathbf{x} \cdot \mathbf{p} \rangle &= \frac{i}{\hbar} \langle [H, \mathbf{x} \cdot \mathbf{p}] \rangle = \frac{i}{\hbar} \langle H \mathbf{x} \cdot \mathbf{p} - \mathbf{x} \cdot \mathbf{p} H \rangle \\ &= \frac{iE}{\hbar} \langle \mathbf{x} \cdot \mathbf{p} - \mathbf{x} \cdot \mathbf{p} \rangle = 0 \end{aligned}$$

  $H = \frac{p^2}{2m} + V(\mathbf{x})$

$$\begin{aligned} 0 &= \left\langle \left[ \frac{p^2}{2m} + V(\mathbf{x}), \mathbf{x} \cdot \mathbf{p} \right] \right\rangle \\ &= \left\langle \left[ \frac{p^2}{2m}, \mathbf{x} \cdot \mathbf{p} \right] + [V(\mathbf{x}), \mathbf{x} \cdot \mathbf{p}] \right\rangle \\ &= \left\langle \frac{1}{2m} [p^2, \mathbf{x}] \cdot \mathbf{p} + \mathbf{x} \cdot [V(\mathbf{x}), \mathbf{p}] \right\rangle \end{aligned}$$

Then as  $[p^2, \mathbf{x}] = \mathbf{p} [p, \mathbf{x}] + [p, \mathbf{x}] \mathbf{p} = -2i\hbar \mathbf{p}$  and  $\mathbf{p} = -i\hbar \nabla_{\mathbf{x}}$ ,

$$0 = \left\langle \frac{-i\hbar}{m} p^2 + i\hbar \mathbf{x} \cdot \nabla_{\mathbf{x}} V(\mathbf{x}) \right\rangle$$

or, the Virial theorem:

$$2 \langle T \rangle = \langle \mathbf{x} \cdot \nabla_x V(\mathbf{x}) \rangle$$

Applied to the harmonic oscillator  $V = 1/2 m\omega^2 x^2$

$$\langle T \rangle = \langle V(\mathbf{x}) \rangle = \frac{1}{2} \hbar \omega \left( n + \frac{1}{2} \right)$$

Applying this to find the RMS excursion of a lattice site in an elemental lattice  $r=1$



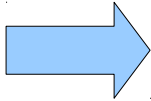
$$\begin{aligned}
\langle s^2 \rangle &= \frac{1}{N} \sum_{n,i} \langle s_{n,i}^2 \rangle \\
&= \frac{1}{N} \left\langle \sum_{n,i} \frac{1}{NM} \sum_{\mathbf{q},s,\mathbf{k},r} Q_s(\mathbf{q}) \epsilon_i^s(\mathbf{q}) e^{i\mathbf{q}\cdot\mathbf{r}_n} Q_r(\mathbf{k}) \epsilon_i^r(\mathbf{k}) e^{i\mathbf{k}\cdot\mathbf{r}_n} \right\rangle \\
&= \frac{1}{N} \left\langle \sum_{n,i} \frac{1}{NM} \sum_{\mathbf{q},s,\mathbf{k},r} Q_s(\mathbf{q}) \epsilon_i^s(\mathbf{q}) e^{i\mathbf{q}\cdot\mathbf{r}_n} Q_r(-\mathbf{k}) \epsilon_i^r(-\mathbf{k}) e^{-i\mathbf{k}\cdot\mathbf{r}_n} \right\rangle \\
&= \frac{1}{N} \left\langle \sum_{n,i} \frac{1}{NM} \sum_{\mathbf{q},s,\mathbf{k},r} Q_s(\mathbf{q}) \epsilon_i^s(\mathbf{q}) e^{i\mathbf{q}\cdot\mathbf{r}_n} Q_r^*(\mathbf{k}) \epsilon_i^{*r}(\mathbf{k}) e^{-i\mathbf{k}\cdot\mathbf{r}_n} \right\rangle \\
&= \frac{1}{NM} \sum_{\mathbf{q},s} \langle |Q_s(\mathbf{q})|^2 \rangle \\
&= \frac{1}{NM} \sum_{\mathbf{q},s} \frac{2}{\omega_s^2(\mathbf{q})} \left\langle \frac{1}{2} \omega_s^2(\mathbf{q}) |Q_s(\mathbf{q})|^2 \right\rangle \\
&= \frac{1}{NM} \sum_{\mathbf{q},s} \frac{2}{\omega_s^2(\mathbf{q})} \frac{1}{2} \hbar \omega_s(\mathbf{q}) \left( n_s(\mathbf{q}) + \frac{1}{2} \right) \\
\langle s^2 \rangle &= \frac{1}{NM} \sum_{\mathbf{q},s} \frac{\hbar}{\omega_s(\mathbf{q})} \left( n_s(\mathbf{q}) + \frac{1}{2} \right)
\end{aligned}$$

However, the integral may be written as a function of  $\omega_s(\mathbf{q})$  only

$$\langle s^2 \rangle = \frac{1}{NM} \sum_{\mathbf{q},s} \frac{\hbar}{\omega_s(\mathbf{q})} \left( \frac{1}{e^{\beta\omega_s(\mathbf{q})} - 1} + \frac{1}{2} \right)$$

It is convenient to introduce a density of phonon states

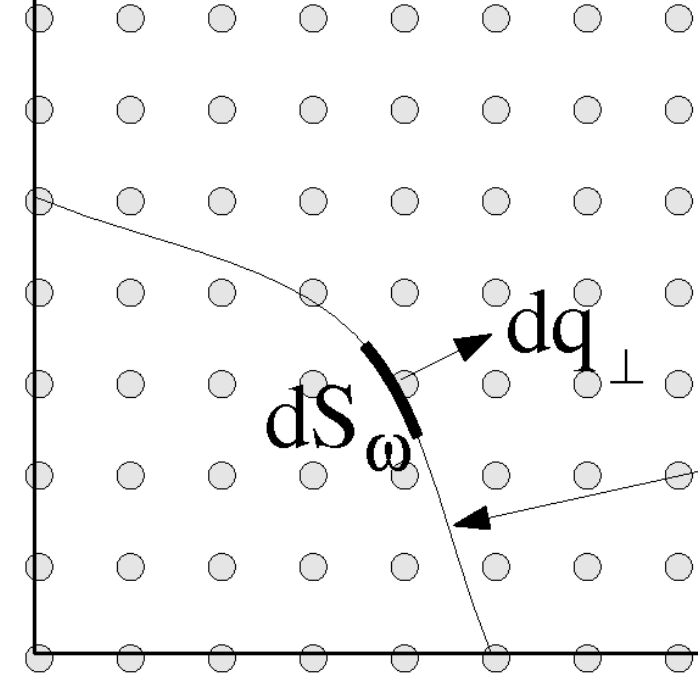
$$Z(\omega) = \frac{1}{N} \sum_{\mathbf{q},s} \delta(\omega - \omega_s(\mathbf{q}))$$

  $\langle s^2 \rangle = \frac{\hbar}{M} \int d\omega Z(\omega) \frac{1}{\omega} \left( n(\omega) + \frac{1}{2} \right)$

# The Phonon Density of States

$$\begin{aligned}
 Z(\omega)d\omega &= \frac{V}{(2\pi)^3} \int_{\omega}^{\omega+d\omega} d^3q \\
 &= \frac{V}{(2\pi)^3} d\omega \sum_s \int d^3q \delta(\omega - \omega_s(\mathbf{q}))
 \end{aligned}$$

$q_y$



$$d^3q = dS_\omega dq_\perp = \frac{dS_\omega d\omega}{\nabla_q \omega_s(\mathbf{q})}$$

Surface

$$\omega = \omega_s(\mathbf{q})$$

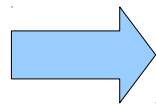
$$d\omega = \nabla_q \omega_s(\mathbf{q}) dq_\perp$$

$q_x$

$S_\omega$  is the surface in  $\mathbf{q}$ -space of constant  $\omega = \omega_s(\mathbf{q})$ ,  
therefor

$$Z(\omega)d\omega = \frac{V}{(2\pi)^3}d\omega \sum_s \int_{\omega=\omega_s(\mathbf{q})} \frac{dS_\omega}{|\nabla_{\mathbf{q}}\omega_s(\mathbf{q})|}$$

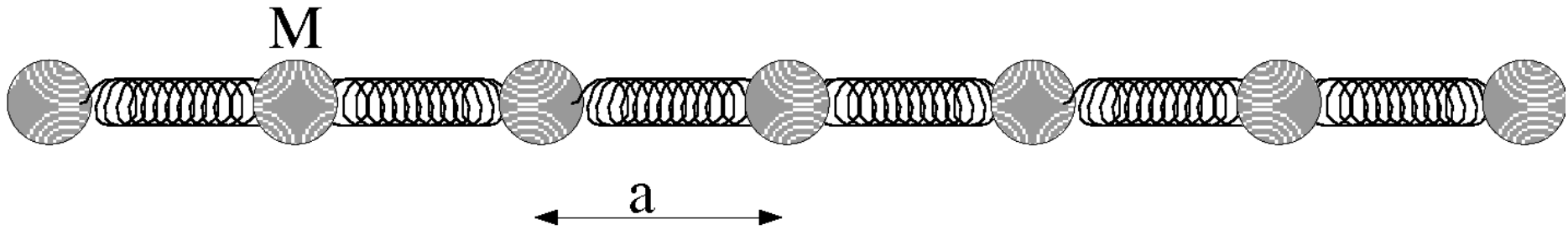
Thus the density of states is high in regions where the dispersion is flat so that  $|\nabla_{\mathbf{q}}\omega_s(\mathbf{q})|$  is small.



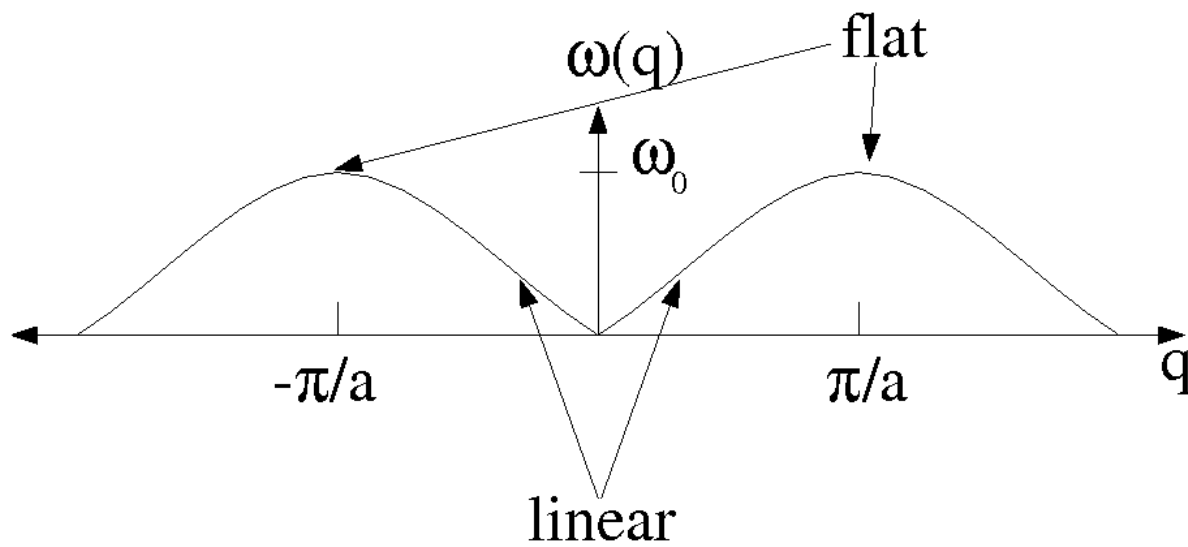
Example: 1-d Harmonic chain



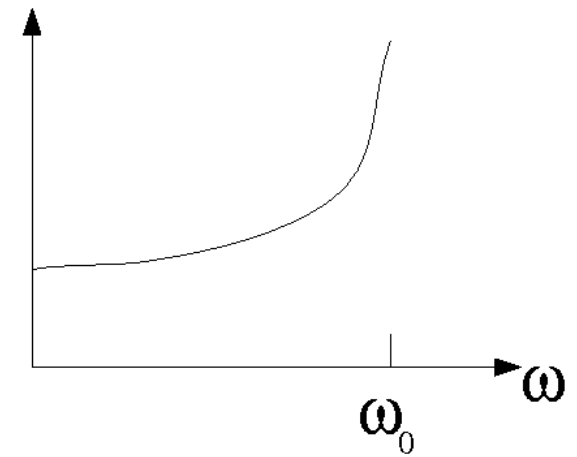
# 1-d Harmonic chain



Dispersion

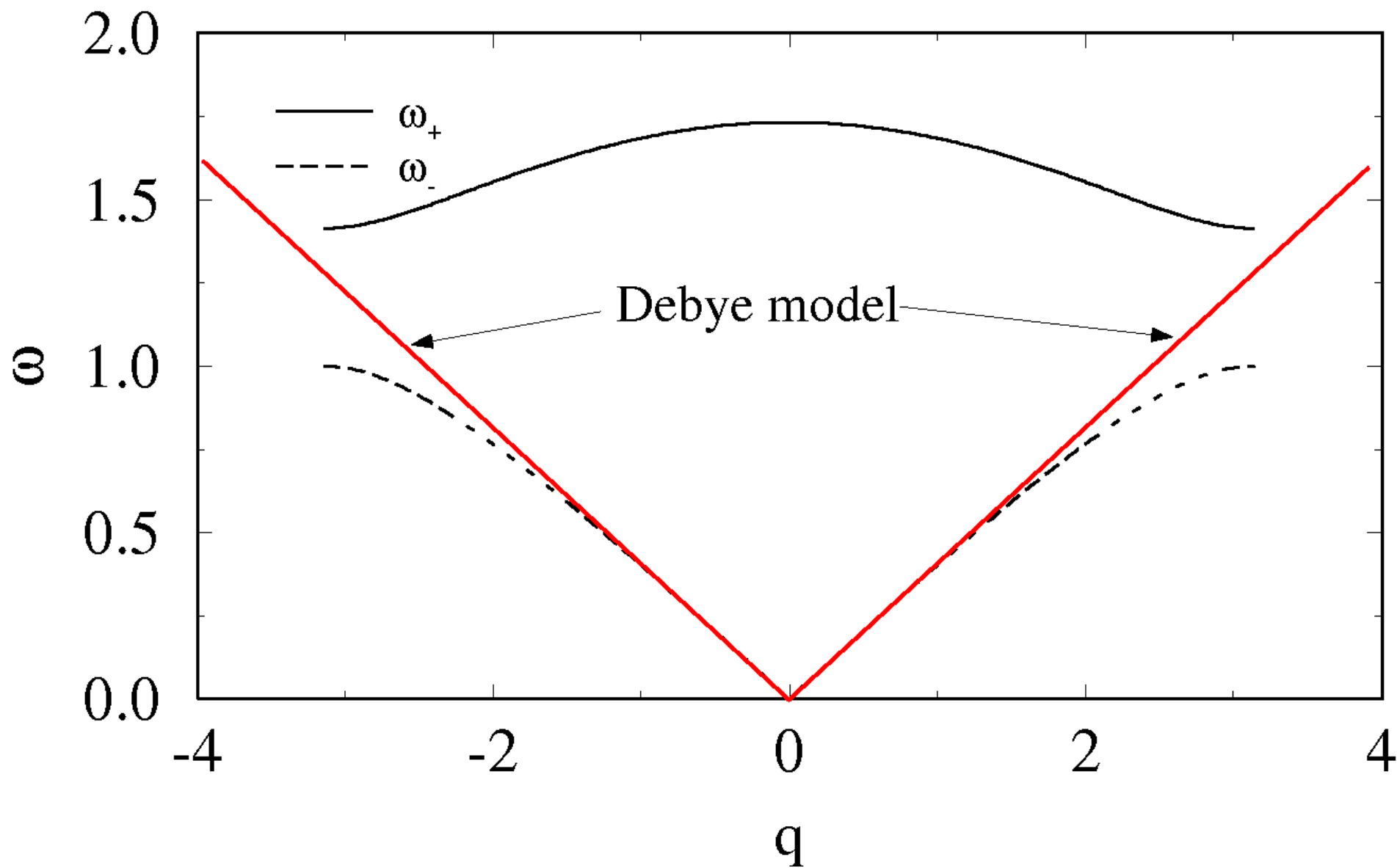


$Z(\omega)$  DOS



# Models of Lattice Dispersion

## The Debye Model



For small  $\omega_s(\mathbf{q}) = c_s|\mathbf{q}|$  for all  $s$  and  $\mathbf{q}$


$\rightarrow \nabla_{\mathbf{q}}\omega_s(\mathbf{q}) = c_s$  for all  $s$  and  $\mathbf{q}$

$$Z(\omega) = \frac{V}{(2\pi)^3} \sum_s \int \frac{dS_\omega}{\nabla_{\mathbf{q}}\omega_s(\mathbf{q})}$$

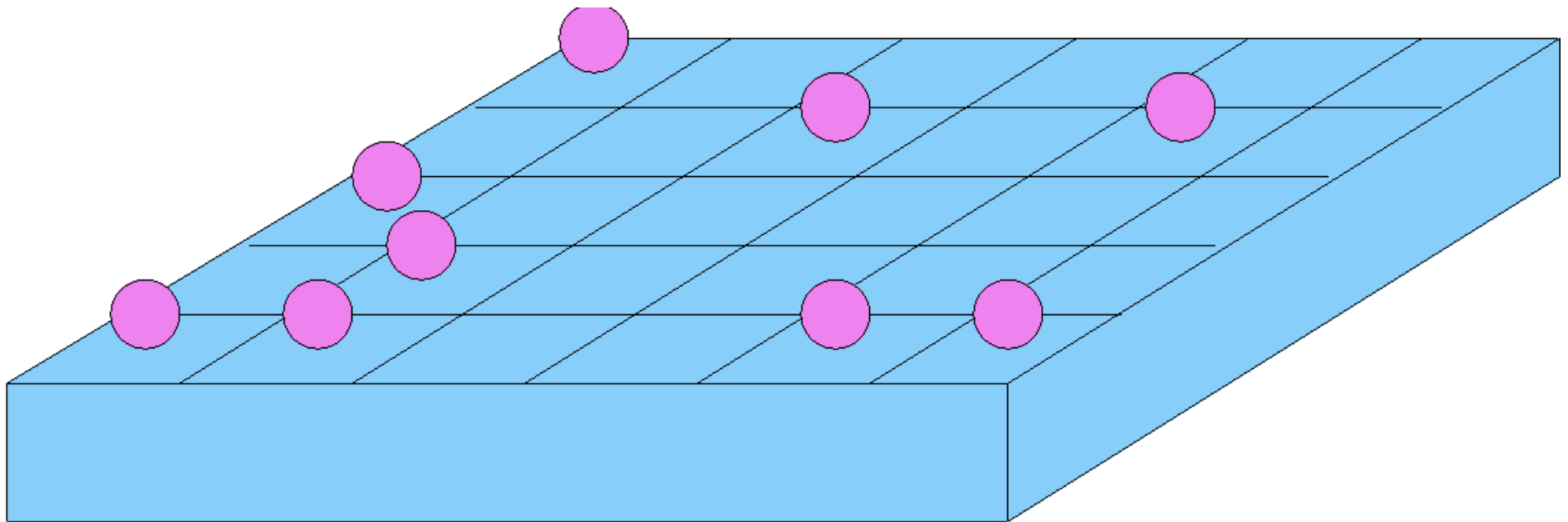
$$= \frac{V}{(2\pi)^3} \sum_s \int \frac{dS_\omega}{c_s}$$

$$S_s(\omega = \omega(\mathbf{q})) = \begin{cases} 2 & \text{for } d = 1 \\ 2\pi q = 2\pi\omega/c & \text{for } d = 2 \\ 4\pi q^2 = 4\pi\omega^2/c^2 & \text{for } d = 3 \end{cases}$$

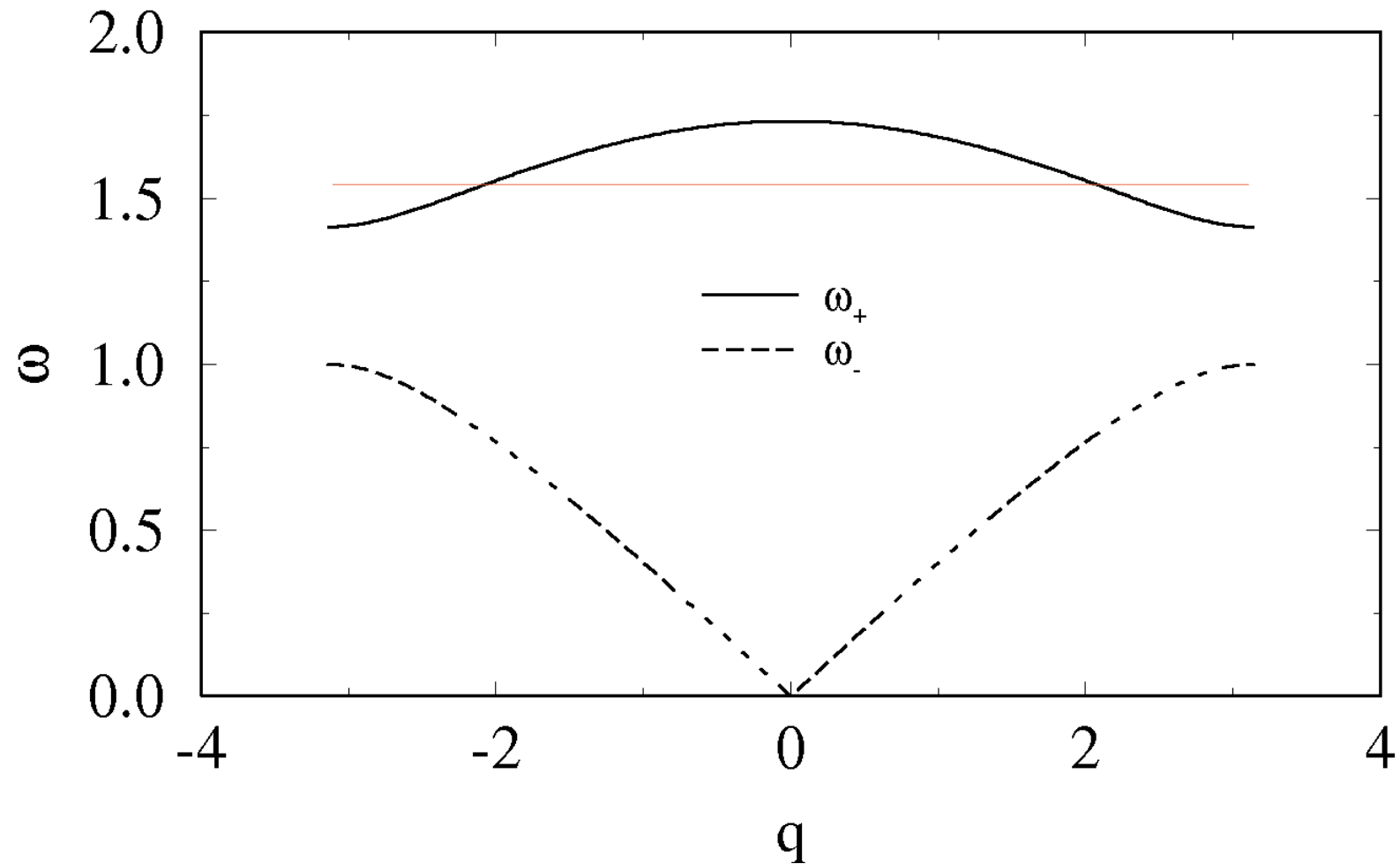
Number of modes =  $d$ :

  $Z(\omega) = \frac{V}{(2\pi)^3} \left\{ \begin{array}{l} 2/c \text{ for } d = 1 \\ 2\pi q = 4\pi\omega/c^2 \text{ for } d = 2 \\ 4\pi q^2 = 12\pi\omega^2/c^3 \text{ for } d = 3 \end{array} \right\} \quad 0 < \omega < \omega_D$

# The Einstein Model



The model is dispersionless  $\omega(q) = \omega_0$ , and the DOS for this system is a delta function  $Z(\omega) = c\delta(\omega - \omega_0)$ .



# Thermodynamics of Crystal Lattices

## Long-Range Order

$$\begin{aligned}\langle s^2 \rangle &= \frac{\hbar}{MN} \sum_{\mathbf{q},s} \frac{1}{\omega_s(\mathbf{q})} \left( \frac{1}{e^{\beta\omega_s(\mathbf{q})} - 1} + \frac{1}{2} \right) \\ &= \frac{\hbar}{2MN} \sum_{\mathbf{q},s} \frac{\sinh(\beta\omega_s(\mathbf{q})/2)}{\omega_s(\mathbf{q}) \cosh(\beta\omega_s(\mathbf{q})/2)}\end{aligned}$$

T  $\rightarrow$  0 limit:

$$\lim_{\beta \rightarrow \infty} \langle s^2 \rangle = \frac{\hbar}{2MN} \lim_{\beta \rightarrow \infty} \sum_{\mathbf{q},s} \frac{1}{\omega_s(\mathbf{q})} \left( \frac{1}{\beta\omega_s(\mathbf{q})} + \frac{1}{2} \right)$$

Low frequency modes are most important → Debye model

$$\lim_{\beta \rightarrow \infty} \langle s^2 \rangle \approx \frac{\hbar}{2MN} \int_0^{\omega_D} d\omega Z(\omega) \frac{1}{\omega} \left( \frac{1}{\beta\omega} + \frac{1}{2} \right)$$

$$Z(\omega) = \frac{V}{(2\pi)^3} \left\{ \begin{array}{l} 2/c \text{ for } d = 1 \\ 2\pi q = 4\pi\omega/c^2 \text{ for } d = 2 \\ 4\pi q^2 = 12\pi\omega^2/c^3 \text{ for } d = 3 \end{array} \right\} \quad 0 < \omega < \omega_D$$

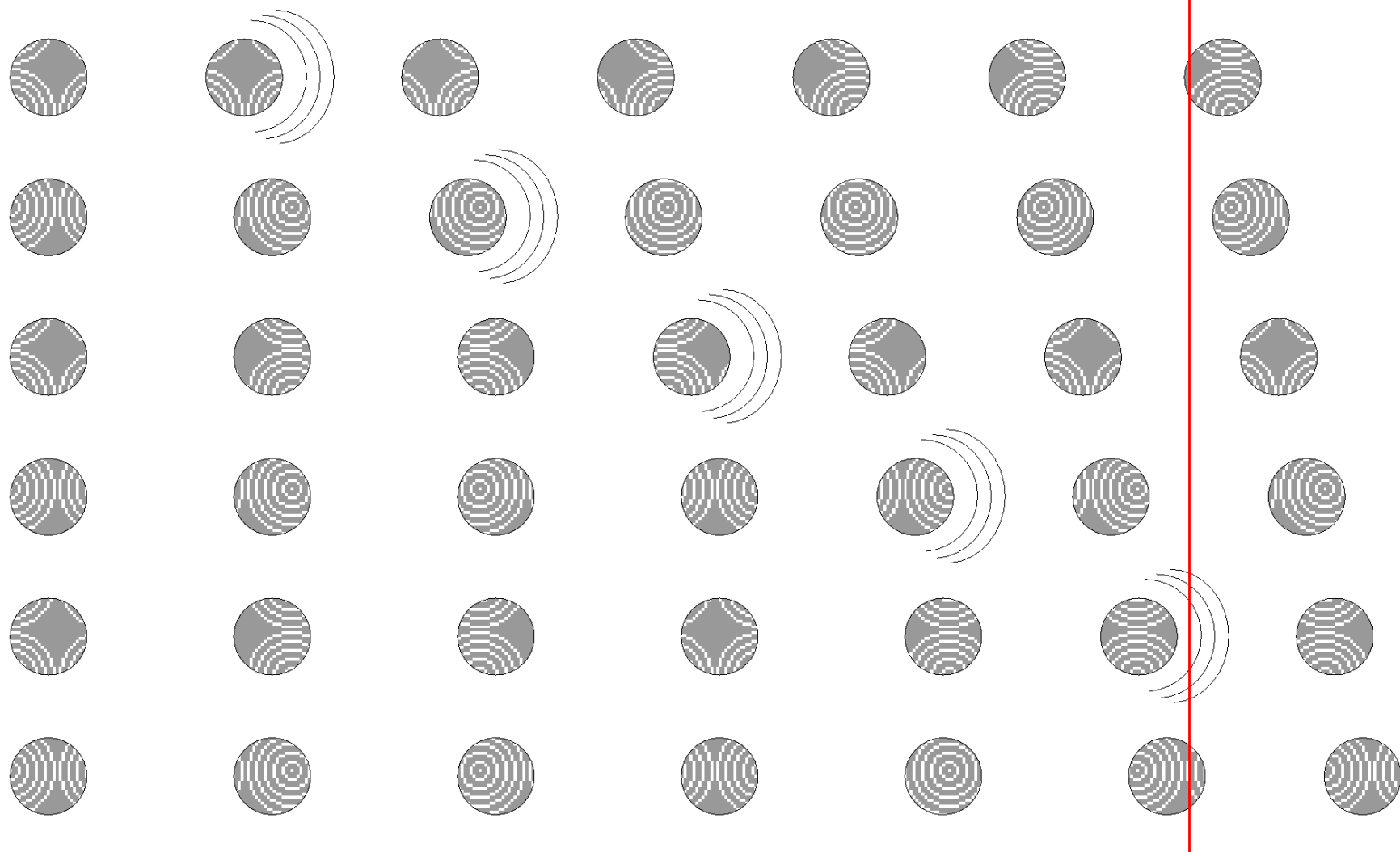
$\langle s^2 \rangle$  for lattices of different dimension

$\langle s^2 \rangle$	$d = 1$	$d = 2$	$d = 3$
$T = 0$	$\infty$	finite	finite
$T \neq 0$	$\infty$	$\infty$	finite

time



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# Thermodynamics

The probability that any state in the subsystem is occupied

$$P(\{n_s(\mathbf{k})\}) \propto e^{-\beta E(\{n_s(\mathbf{k})\})}$$

$$\begin{aligned} \mathcal{Z} &= \sum_{\{n_s(\mathbf{k})\}} e^{-\beta E(\{n_s(\mathbf{k})\})} \\ &= \sum_{\{n_s(\mathbf{k})\}} e^{-\beta \sum_{\mathbf{k},s} \hbar \omega_s(\mathbf{k}) (n_s(\mathbf{k}) + \frac{1}{2})} \\ &= \prod_{s,\mathbf{k}} \mathcal{Z}_s(\mathbf{k}) \end{aligned}$$

$$\begin{aligned}
\mathcal{Z}_s(\mathbf{k}) &= \sum_n e^{-\beta\hbar\omega_s(\mathbf{k})(n_s(\mathbf{k})+\frac{1}{2})} \\
&= e^{-\beta\hbar\omega_s(\mathbf{k})/2} \sum_n e^{-\beta\hbar\omega_s(\mathbf{k})(n_s(\mathbf{k}))} \\
&= \frac{e^{-\beta\hbar\omega_s(\mathbf{k})/2}}{1 - e^{-\beta\hbar\omega_s(\mathbf{k})}} \\
&= \frac{1}{2 \sinh(\beta\omega_s(\mathbf{k})/2)}
\end{aligned}$$

$$\mathcal{F} = -k_B T \ln(\mathcal{Z}) = k_B T \sum_{\mathbf{k},s} \ln(2 \sinh(\beta\hbar\omega_s(\mathbf{k})/2))$$

Since  $dE = TdS - PdV$  and  $dF = TdS - PdV - TdS - SdT$ , the entropy is

$$S = - \left( \frac{\partial \mathcal{F}}{\partial T} \right)_V$$

and system energy is then given by

$$\mathcal{E} = \mathcal{F} + TS = \mathcal{F} - T \left( \frac{\partial \mathcal{F}}{\partial T} \right)_V$$

where constant volume  $V$  is guaranteed by the harmonic approximation (since  $\langle s \rangle = 0$ )

$$\mathcal{E} = \sum_{\mathbf{k}, s} \frac{1}{2} \hbar \omega_s(\mathbf{k}) \coth(\beta \omega_s(\mathbf{k})/2)$$

The specific heat is then given by

$$C = \left( \frac{d\mathcal{E}}{dT} \right)_V = k_B \sum_{\mathbf{k},s} (\beta \hbar \omega_s(\mathbf{k}))^2 \operatorname{csch}^2 (\beta \omega_s(\mathbf{k})/2)$$

where  $\operatorname{csch}(x) = 1/\sinh(x)$

Consider the specific heat of our 3-dimensional Debye model.

$$\begin{aligned} C &= k_B \int_0^{\omega_D} d\omega \mathcal{Z}(\omega) (\beta \omega/2)^2 \operatorname{csch}^2 (\beta \omega/2) \\ &= k_B \int_0^{\omega_D} d\omega \left( \frac{12V \pi \omega^2}{(2\pi c)^3} \right) (\beta \omega/2)^2 \operatorname{csch}^2 (\beta \omega/2) \end{aligned}$$

Debye frequency  $\omega_D$  is determined by the requirement that

$$3rN = \int_0^{\omega_D} d\omega \mathcal{Z}(\omega) = \int_0^{\omega_D} d\omega \left( \frac{12V\pi\omega^2}{(2\pi c)^3} \right)$$

or  $V/(2\pi c)^3 = 3rN/(4\omega_D^3)$ .



At high temperatures  
 $\beta\hbar\omega_D/2 \ll 1$

$$C \approx k_B \int_0^{\omega_D} d\omega \mathcal{Z}(\omega) = 3Nr k_B$$



At low temperatures

$$\beta \hbar \omega_D / 2 \gg 1$$

$$\operatorname{csch}^2(\beta \omega / 2) \approx 2e^{-\beta \omega / 2}$$

$$C \approx 12\pi k_B \frac{3rN}{4\omega_D^2} \int_0^\infty d\omega \omega^2 \left( \frac{\beta \hbar \omega}{2} \right)^2 2e^{-\beta \omega_s(\mathbf{k})/2}$$

$$C \approx \frac{9k_B r N \pi}{2} \left( \frac{1}{\omega_D \beta \hbar} \right)^3 \int_0^\infty dx x^4 e^{-x}$$

$$\approx \frac{9k_B r N \pi}{2} \left( \frac{1}{\omega_D \beta \hbar} \right)^3 24$$

Then, if we identify the Debye temperature  $\theta_D = \hbar \omega_D / k_B$ , we get

$$C \approx 96\pi r N k_B \left( \frac{T}{\theta_D} \right)^3$$

## Einstein Model:

$$C = pnk_B \frac{(\hbar\omega_E/k_B T)^2 e^{\hbar\omega_E/k_B T}}{(e^{\hbar\omega_E/k_B T} - 1)^2}$$



At low temperatures  
 $T \ll \Theta_E$

**Constant**



At high temperatures  
 $T \gg \Theta_E$

**Exponential**

Were we have used the Debye temperature

$$\Theta_E = \hbar\omega_e/k_B$$

# Specific Heat

